

DOING PHYSICS WITH PYTHON

QUANTUM MECHANICS

BLACKBODY RADIATION SUN, RED STAR, BLUE STAR

Ian Cooper

matlabvisualphysics@gmail.com

DOWNLOAD DIRECTORY FOR MATLAB SCRIPTS

qmSun.py

[GitHub](#)

[Google Drive](#)

BLACKBODY RADIATION

The wave nature of electromagnetic radiation is demonstrated by interference phenomena. However, electromagnetic radiation also has a particle nature. For example, to account for the observations of the radiation emitted from hot objects, it is necessary to use a

particle model, where the radiation is considered to be a stream of particles called **photons**. The energy of a photon, E is

$$(1) \quad E = h f$$

The electromagnetic energy emitted from an object's surface is called **thermal radiation** and is due a decrease in the internal energy of the object. This radiation consists of a continuous spectrum of frequencies extending over a wide range. Objects at room temperature emit mainly infrared and it is not until the temperature reaches about 800 K and above those objects glows visibly.

A **blackbody** is an object that completely absorbs all electromagnetic radiation falling on its surface at any temperature. It can be thought of as a perfect absorber and emitter of radiation. The power emitted from a blackbody, P is given by the **Stefan-Boltzmann law** and it depends only on the surface area of the emitter, A and its surface temperature, T

$$(2) \quad P = A \sigma T^4$$

A more general form of equation 2 is

$$(2) \quad P = \varepsilon A \sigma T^4$$

where ε is the **emissivity** of the object. For a blackbody, $\varepsilon = 1$. When $\varepsilon < 1$ the object is called a **graybody** and the object is not a perfect emitter and absorber.

The amount of radiation emitted by a blackbody is given by **Planck's radiation law** and is expressed in terms of the **spectral exitance** for **wavelength** or **frequency** R_λ or R_f respectively

$$(4) \quad R_\lambda = \frac{2\pi h c^2}{\lambda^5} \frac{1}{\exp\left(\frac{hc}{k_B T \lambda}\right) - 1} \quad [\text{W.m}^{-2}.\text{m}^{-1}]$$

or

$$(5) \quad R_f = \frac{2\pi h f^3}{c^2} \frac{1}{\exp\left(\frac{hf}{k_B T}\right) - 1} \quad [\text{W.m}^{-2}.\text{s}^{-1}]$$

In the literature, many different terms and symbols are used for the spectral exitance. Sometimes the terms and the units given are wrong or misleading.

The **power radiated per unit surface of a blackbody**, P_A within a wavelength interval or bandwidth, (λ_1, λ_2) or frequency interval or bandwidth (f_1, f_2) are given by equations 6 and 7

$$(6) \quad P_A = \int_{\lambda_1}^{\lambda_2} R_\lambda d\lambda = \int_{\lambda_1}^{\lambda_2} \left(\frac{2\pi h c^2}{\lambda^5} \frac{1}{\exp\left(\frac{hc}{k_b T}\right) - 1} \right) d\lambda \quad [\text{W.m}^{-2}]$$

and

$$(7) \quad P_A = \int_{f_1}^{f_2} R_f df = \int_{f_1}^{f_2} \left(\frac{2\pi h f^3}{c^2} \frac{1}{\exp\left(\frac{hf}{k_b T}\right) - 1} \right) df \quad [\text{W.m}^{-2}]$$

The equations 6 and 7 give the Stefan-Boltzmann law (equation 2) when the bandwidths extend from 0 to ∞ .

Wien's Displacement law states that the wavelength λ_{peak} corresponding to the peak of the spectral exitance given by equation 4 is inversely proportional to the temperature of the blackbody and the frequency f_{peak} for the spectral exitance peak frequency given by equation 5 is proportional to the temperature

$$(8) \quad \lambda_{peak} = \frac{b_\lambda}{T} \quad f_{peak} = b_f T$$

The peaks in equations 4 and 5 occur in different parts of the electromagnetic spectrum and so

$$(9) \quad f_{peak} \neq \frac{c}{\lambda_{peak}}$$

The Wien's Displacement law explains why long wave radiation dominates more and more in the spectrum of the radiation emitted by an object as its temperature is lowered.

When classical theories were used to derive an expression for the spectral exitances R_λ and R_f , the power emitted by a blackbody diverged to infinity as the wavelength became shorter and shorter. This is known as the **ultraviolet catastrophe**. In 1901 Max Planck proposed a new radical idea that was completely alien to classical notions, electromagnetic energy is **quantized**. Planck was able to derive the equations 4 and 5 for blackbody emission and these equations are in complete agreement with experimental measurements. The assumption that the energy of a system varies in a continuous manner, i.e., it can take any arbitrary close consecutive values fails. Energy can only exist in integer multiples of the lowest amount or quantum, hf . ***This step marked the very beginning of modern quantum theory.***

A summary of the physical quantities, units and values of constants used in the description of the radiation from a hot object.

Variable	Interpretation	Value	Unit
E	energy of photon		J, eV
h	Planck's constant	6.62608×10^{-34}	J.s
c	speed of electromagnetic radiation	3.00×10^8	m.s^{-1}
f	frequency of electromagnetic radiation		Hz
λ	wavelength of electromagnetic radiation		
T	surface temperature of object		K
A	surface area of object		m^2
σ	Stefan-Boltzmann constant	5.6696×10^{-8}	$\text{W.m}^{-2}.\text{K}^{-4}$
P	power emitted from hot object		W
ε	emissivity of object's surface		
R_λ	spectral exitance: power radiated per unit area per unit wavelength interval		$(\text{W.m}^{-2}).\text{m}^{-1}$
R_f	spectral exitance: power radiated per unit area per unit frequency interval		$(\text{W.m}^{-2}).\text{s}^{-1}$
k_B	Boltzmann constant	1.38066×10^{-23}	J.K^{-1}
b_λ	Wien constant: wavelength	2.898×10^{-3}	m.K
b_f	Wien constant: frequency	$2.83 k_B T / h$	$\text{K}^{-1}.\text{s}^{-1}$

λ_{peak}	wavelength of peak in solar spectrum	5.0225×10^{-7}	m
R_S	radius of the Sun	6.96×10^8	m
R_E	radius of the Earth	6.96×10^6	m
R_{SE}	Sun-Earth radius	6.96×10^{11}	m
I_0	Solar constant	1.36×10^3	W.m^{-2}
α	Albedo of Earth's surface	0.30	

SIMULATION: THE SUN AND THE EARTH AS BLACKBODIES

The Sun can be considered as a blackbody, and the total power output of the Sun P_S can be estimated by using the Stefan-Boltzmann law, equation 2, and by finding the area under the curves for R_λ and R_f using equations 6 and 7. From observations on the Sun, the peak in the electromagnetic radiation emitted has a wavelength, $\lambda_{peak} = 502.25 \text{ nm}$ (green). The temperature of the Sun's surface (photosphere) can be estimated from the Wien displacement law, equation 8.

The distance from the Sun to the Earth, R_{SE} can be used to estimate of the surface temperature of the Earth T_E if there was no atmosphere. The intensity of the Sun's radiation reaching the top of the atmosphere, I_0 is known as the **solar constant**

$$(10) \quad I_0 = \frac{P_S}{4\pi R_{SE}^2}$$

The power absorbed by the Earth, P_{Eabs} is

$$(11) \quad P_{Eabs} = (1 - \alpha)\pi R_E^2 I_0$$

where α is the albedo (the reflectivity of the Earth's surface).

Assuming the Earth behaves as a blackbody then the power of the radiation emitted from the Earth, P_{Erad} is

$$(12) \quad P_{Erad} = 4\pi R_E^2 \sigma T_E^4$$

It is known that the Earth's surface temperature has remained relatively constant over many centuries, so that the power absorbed and the power emitted are equal, so the Earth's equilibrium temperature T_E is

$$(13) \quad T_E = \left(\frac{(1 - \alpha)I_0}{4\sigma} \right)^{0.25}$$

Simulation using `qmSun.py`

Console output

Sun: temperature of photosphere, $T_S = 5770$ K

Peak in Solar Spectrum

Theory: Wavelength at peak in spectral exitance

$$wL_{\text{peak}} = 5.02e-07 \text{ m}$$

Graph: Wavelength at peak in spectral exitance

$$wL_{\text{peak}} = 5.04e-07 \text{ m}$$

Theory: Frequency at peak in spectral exitance

$$f_{\text{peak}} = 3.39e+14 \text{ Hz}$$

Graph: Frequency at peak in spectral exitance

$$f_{\text{peak}} = 3.40e+14 \text{ Hz}$$

Total Solar Power Output

$$P_{\text{Stefan_Boltzmann}} = 3.79e+26 \text{ W}$$

$$P_{wL} = 3.77e+26 \text{ W}$$

$$P_f = 3.79e+26 \text{ W}$$

IR / visible / UV

$$P_{\text{IR}} = 1.92e+26 \text{ W} \quad \text{percentage } 50.95$$

$$P_{\text{vis}} = 1.39e+26 \text{ W} \quad \text{percentage } 36.82$$

$$\text{UV} = 4.61e+25 \text{ W} \quad \text{percentage } 12.23$$

Sun - Earth

Theory: Solar constant $I_O = 1.360e+03 \text{ W/m}^2$

Computed: Solar constant $I_E = 1.34e+03 \text{ W/m}^2$

Surface temperature of the Earth, $T_E = 254 \text{ K} = -19 \text{ deg C}$

Execution time: 41.06

Without our atmosphere, the Earth's temperature would be $\sim 19^\circ\text{C}$

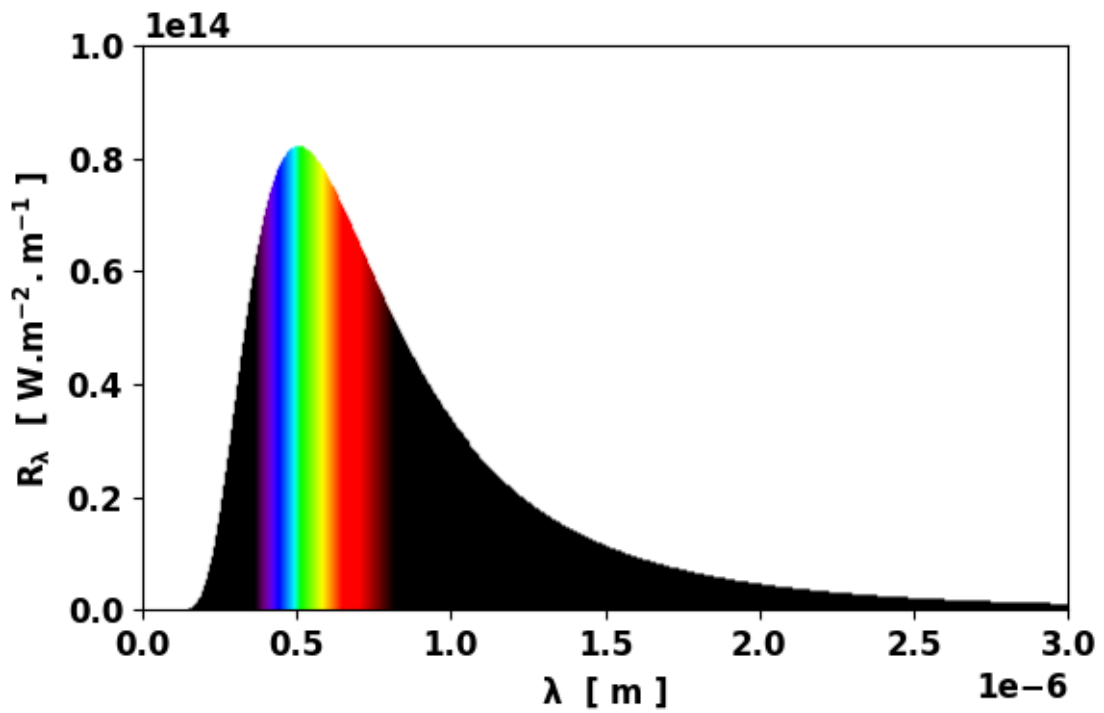


Fig. 1. Spectral exitance curve $T = 5770\text{ K}$

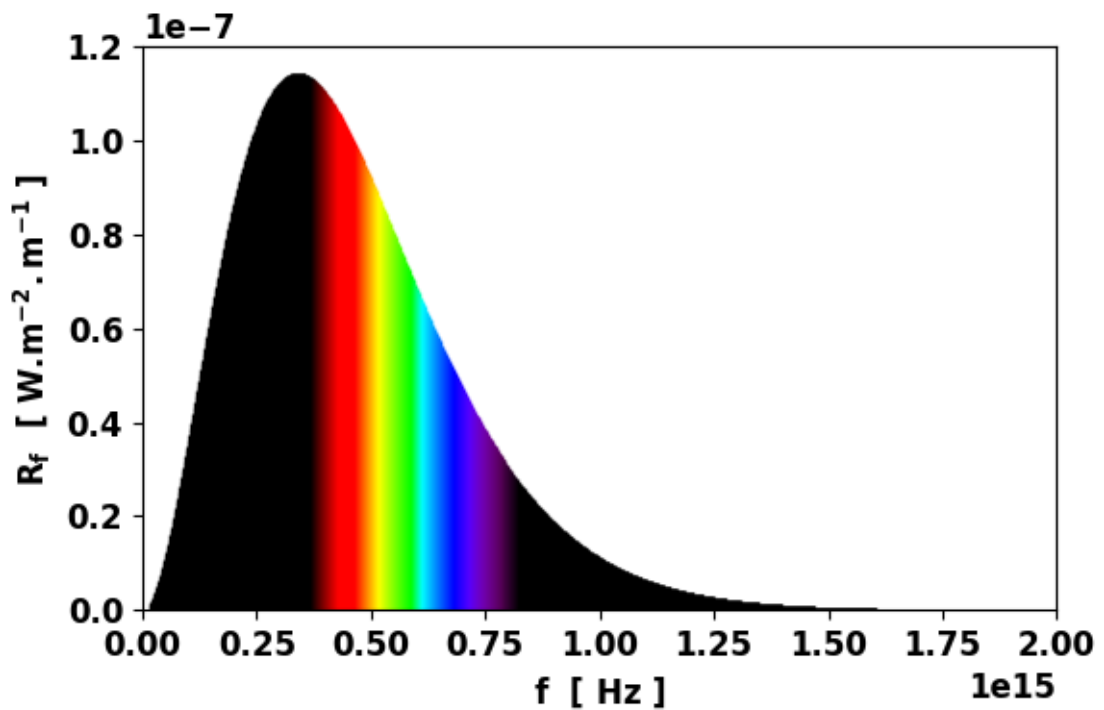


Fig. 2. Spectral exitance curve $T = 5770\text{ K}$

SIMULATION: STAR TEMPERATURES



Stars approximate blackbody radiators and their visible color depends upon the temperature of the radiator. Our Sun with a photosphere temperature ~ 6000 K is a yellow-white star. The curves below are for a **blue** star (7000 K) and a **red** star (4000 K).

BLUE STAR

Total Solar Power Output

$$P_{\text{Stefan_Boltzmann}} = 8.22e+26 \text{ W}$$

$$P_{\text{wL}} = 8.19e+26 \text{ W}$$

$$P_{\text{f}} = 8.21e+26 \text{ W}$$

IR / visible / UV

$$P_{\text{IR}} = 3.09e+26 \text{ W} \quad \text{percentage } 37.72$$

$$P_{\text{vis}} = 3.23e+26 \text{ W} \quad \text{percentage } 39.42$$

$$\text{UV} = 1.87e+26 \text{ W} \quad \text{percentage } 22.86$$

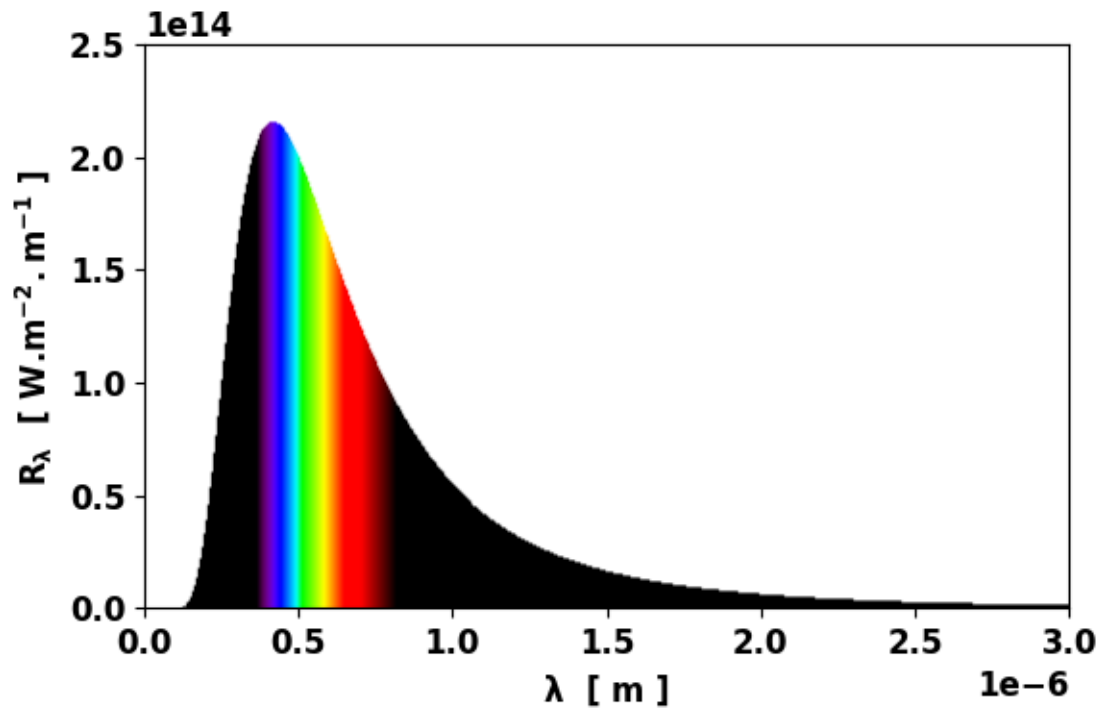


Fig. 3. Spectral exitance curve $T = 7000 \text{ K}$

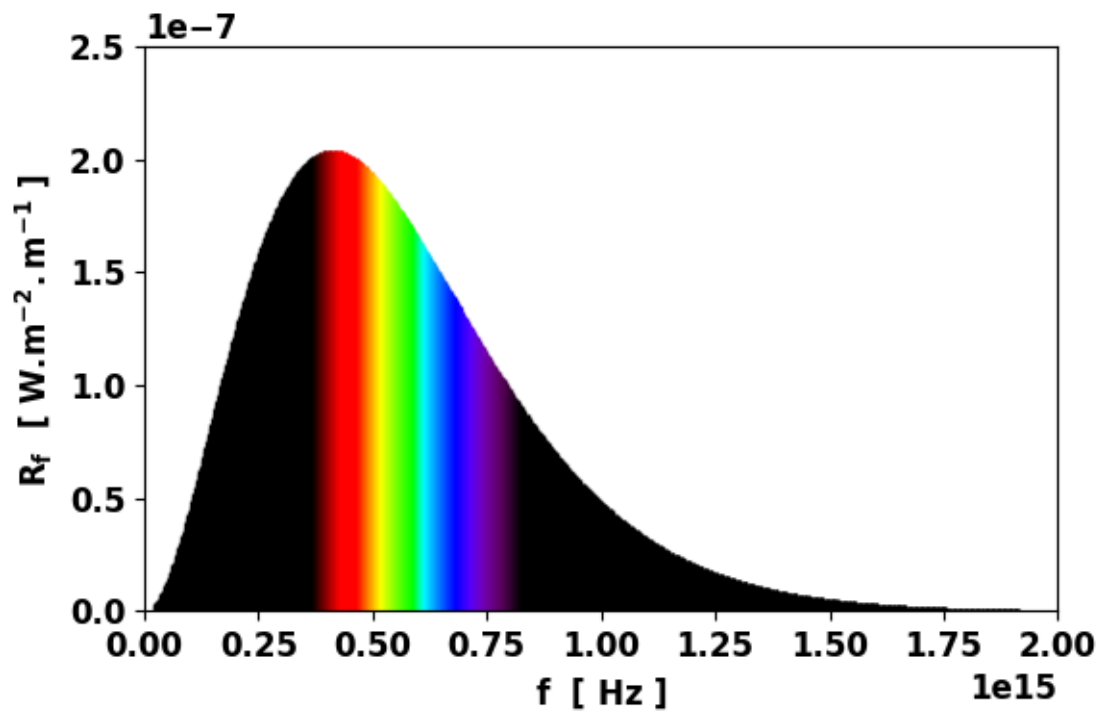


Fig. 4. Spectral exitance curve $T = 7000 \text{ K}$

RED STAR

Total Solar Power Output

$$P_{\text{Stefan_Boltzmann}} = 8.76e+25 \text{ W}$$

$$P_{\text{wL}} = 8.64e+25 \text{ W}$$

$$P_{\text{f}} = 8.76e+25 \text{ W}$$

IR / visible / UV

$$P_{\text{IR}} = 6.64e+25 \text{ W} \text{ percentage } 76.88$$

$$P_{\text{vis}} = 1.82e+25 \text{ W} \text{ percentage } 21.12$$

$$\text{UV} = 1.73e+24 \text{ W} \text{ percentage } 2.00$$

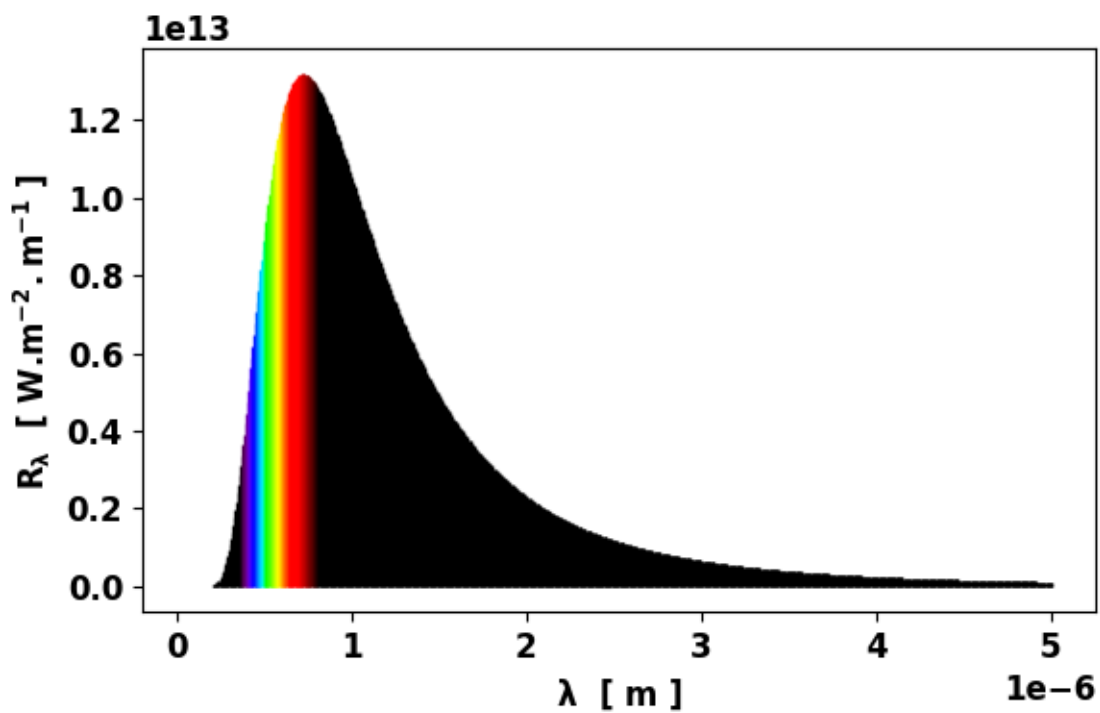


Fig. 5. Spectral exitance curve $T = 4000 \text{ K}$

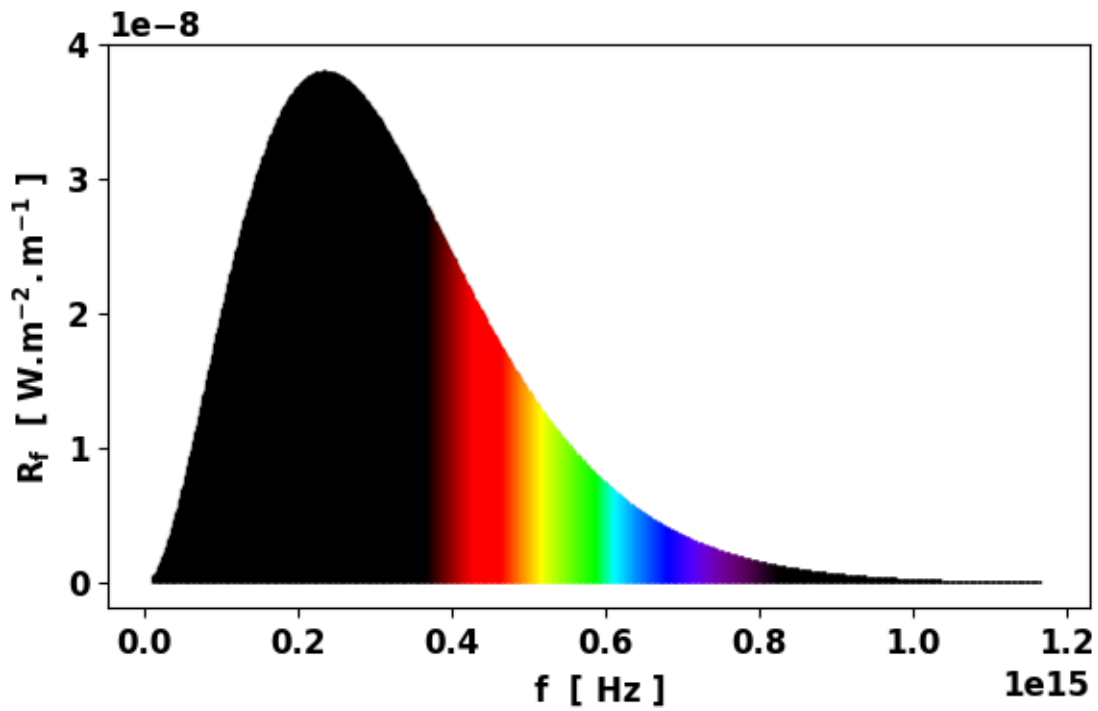


Fig. 6. Spectral exitance curve $T = 4000\text{ K}$

Comparison of radiation emitted by the stars.

	Red Star	Sun yellow- white	Blue Star
$T_{\text{star}}\text{ [K]}$	4000	5770	7000
$P\text{ [W]}$	1×10^{26}	4×10^{26}	8×10^{26}
% IR	77	51	38
% visible	21	37	39
% UV	2	12	23